EUI High Resolution Observations of **Decayless Oscillations**

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• Decayless oscillations are routinely observed with AIA and now also with HRIEUV. Co-observations EUI & AIA detect the same decayless oscillations, confirming that they are of solar origin.

- Decayless oscillations are different from flare-induced oscillations as they do not dampen during their lifetime and they are independent of flares. In fact the driver is unknown.
- •These oscillations have been interpreted as kink-waves and the ratio between the loop length and the period tends to be in the order of 1 Mm/s, suggestive of the coronal Alfvén speed.



HRIEUV revealed the smallest decayless oscillations so far: loops of less than 5 Mm, oscillating with a period as fast as 14s.



Decayless oscillations in short loops have relatively long periods, with a ratio loop length/period of only ~100km/s.



Meta-analysis of 290 oscillations (AIA and EUI) by 10 authors. Higher energy fluxes are generated from higher-frequency oscillations. The total energy generated by transverse oscillations ranges from about 10^20 to 10^25 erg, (cfr energies for nanoflare 10^24 to 10^27 erg). The respective slope results imply that high-frequency oscillations could provide the dominant contribution to total coronal heating generated by decayless transverse oscillations.

To be explored: what drives decayless oscillations? why do we have two branches with different ratio loop length/period? are there enough decayless oscillations in different coronal zones

oscillations in different coronal zones for coronal heating?

Find out more:

- Zhong et al 2022MNRAS.516.5989Z
- Mandal et al 2022A&A...666L...2M
- Petrova et al 2023ApJ...946...36P
 Lim et al 2023ApJ...952L..15L
- Lim et al 2023ApJ...952L..15L
 Zhong et al 2023NatCo..14.5298Z
- Shrivastav et al 2023NatCo...14.52962

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